

1. The key to this problem is to realize that the  $\beta$ s in the expression for  $\gamma$  have a time derivative in them. So an argument of this type is valid

$$\int_{t_1}^{t_2} \sqrt{1 - \beta^2} dt = \int_{t_1}^{t_2} \sqrt{dt^2 - (1/c^2)(dx^2 + dy^2 + dz^2)}.$$

But the invariance of the space-time interval shows that under a general Lorentz transformation

$$\int_{t_1}^{t_2} \sqrt{dt^2 - (1/c^2)(dx^2 + dy^2 + dz^2)} = \int_{t'_1}^{t'_2} \sqrt{dt'^2 - (1/c^2)(dx'^2 + dy'^2 + dz'^2)}$$

Therefore

$$\int_{t_1}^{t_2} \sqrt{1 - \beta^2} dt = \int_{t'_1}^{t'_2} \sqrt{1 - \beta'^2} dt'.$$

The square root differential expression is well defined because the four-velocity of any massive particle is always time-like.

A more formal and precise argument is the following. Divide the closed interval  $[t_1, t_2]$  into  $N$  equal sub-intervals of duration  $\Delta t = (t_2 - t_1)/N$  labeled by the index  $i$ :  $I_i = [t_1 + (i-1)\Delta t, t_1 + i\Delta t]$ . By the general Lorentz transformation between frames we may establish the coordinates of the space-time events  $c(t_1), \vec{x}(t_1)$  and  $c(t_1 + i\Delta t), \vec{x}(t_1 + i\Delta t)$  in the prime frame. Call the coordinates in the  $K'$  frame  $c(t'_0), \vec{x}'_0$  and  $c(t'_i), \vec{x}'_i$  respectively. Recall the mean value theorem from calculus

$$\int_{t_1}^{t_2} \sqrt{1 - \beta^2} dt = \sum_{i=1}^N \int_{I_i} \sqrt{1 - \beta^2} dt = \sum_{i=1}^N \sqrt{1 - \beta^2(T_i)} \Delta t$$

for some  $T_i \in I_i$ . In the limit  $N \rightarrow \infty$ , the intervals become infinitesimals and

$$\vec{\beta}(T_i) \rightarrow \frac{\Delta \vec{x}_i}{\Delta t} \equiv \frac{\vec{x}_i - \vec{x}_{i-1}}{\Delta t}.$$

This means

$$\int_{t_1}^{t_2} \sqrt{1 - \beta^2} dt = \lim_{N \rightarrow \infty} \sum_{i=1}^N \sqrt{\Delta t^2 - (1/c^2)|\Delta \vec{x}_i|^2}$$

The Lorentz transformations are linear, and so the differentials  $c\Delta t$  and  $\Delta \vec{x}$  transform in the same way as  $ct$  and  $\vec{x}$ . Therefore, by the invariance of the space-time interval under Lorentz transformations,

$$\int_{t_1}^{t_2} \sqrt{1 - \beta^2} dt = \lim_{N \rightarrow \infty} \sum_{i=1}^N \sqrt{\Delta t_i'^2 - (1/c^2)|\Delta \vec{x}'_i|^2} = \int_{t'_1}^{t'_2} \sqrt{1 - \beta'^2} dt',$$

because the durations of all the intervals in the prime frame approach zero as  $N$  increases without bound. It should be noted that  $\Delta t'_i$  is NOT necessarily constant as  $i$  changes when there is acceleration in the orbit.

2. By the relativistic Lorentz Force equation

$$\vec{v} \cdot \frac{d(\gamma m \vec{v})}{d\tau} = \vec{v} \cdot q(\vec{E} + \vec{v} \times \vec{B}) = q \vec{v} \cdot \vec{E}$$

Now

$$\frac{d\gamma}{dt} = \frac{d}{dt} \frac{1}{\sqrt{1 - |\vec{v}|^2 / c^2}} = \frac{\vec{v} \cdot (d\vec{v} / dt)}{c^2 (1 - \beta^2)^{3/2}} = \frac{\gamma^3}{c^2} \vec{v} \cdot (d\vec{v} / dt)$$

and

$$\begin{aligned} q \vec{v} \cdot \vec{E} &= \gamma \vec{v} \cdot \frac{d(m \vec{v})}{dt} + \frac{m |\vec{v}|^2}{c^2} \gamma^3 \vec{v} \cdot (d\vec{v} / dt) \\ &= \gamma m (1 + \beta^2 \gamma^2) \vec{v} \cdot (d\vec{v} / dt) = \gamma^3 m \vec{v} \cdot (d\vec{v} / dt). \end{aligned}$$

So

$$\frac{d\gamma}{dt} = \frac{q \vec{v} \cdot \vec{E}}{mc^2}.$$

Finally

$$\begin{aligned} \frac{d(\gamma m v_x)}{d\tau} &= q E_x u^0 + 0 + q c B_z u^2 - q c B_y u^3 \\ \gamma \frac{d(\gamma m v_x)}{dt} &= q E_x \gamma + q v_y B_z \gamma - q v_z B_y \gamma \\ \frac{d(\gamma m v_x)}{dt} &= q (\vec{E} + \vec{v} \times \vec{B})_x \\ \frac{d(\gamma m v_y)}{d\tau} &= q E_y u^0 - q c B_z u^1 + 0 + q c B_x u^3 \\ \gamma \frac{d(\gamma m v_y)}{dt} &= q E_y \gamma - q v_x B_z \gamma + q v_z B_x \gamma \\ \frac{d(\gamma m v_y)}{dt} &= q (\vec{E} + \vec{v} \times \vec{B})_y \\ \frac{d(\gamma m v_z)}{d\tau} &= q E_z u^0 + q c B_y u^1 - q c B_x u^2 + 0 \\ \gamma \frac{d(\gamma m v_z)}{dt} &= q E_z \gamma + q v_x B_y \gamma - q v_y B_x \gamma \\ \frac{d(\gamma m v_z)}{dt} &= q (\vec{E} + \vec{v} \times \vec{B})_z \end{aligned}$$

### 3. My solution

$$\begin{aligned}\frac{d\gamma}{dt} = 0 &\rightarrow \frac{d|\vec{v}|}{dt} = 0 \quad |\vec{v}| = \text{const} = v_0 \\ \frac{dv_x}{dt} &= \frac{qB}{\gamma m} v_y \quad \frac{dv_y}{dt} = -\frac{qB}{\gamma m} v_x \\ \frac{d^2 v_x}{dt^2} + \Omega_c^2 v_x &= 0 \quad \frac{d^2 v_y}{dt^2} + \Omega_c^2 v_y = 0 \\ v_x(t) &= A \cos(\Omega_c t + \delta) \quad v_y(t) = -A \sin(\Omega_c t + \delta) \\ |\vec{v}| &= A \rightarrow A = v_0 \\ x(t) &= x_c + \frac{v_0}{\Omega_c} \sin(\Omega_c t + \delta) \quad y(t) = y_c + \frac{v_0}{\Omega_c} \cos(\Omega_c t + \delta) \\ r &= \frac{v_0}{\Omega_c} = \frac{\beta c}{qB/\gamma m}\end{aligned}$$

It turns out that Serkan Golge doesn't like me to invoke "unphysical" things like time derivatives of acceleration. For his benefit one can also get it like this

$$\begin{aligned}v_x(t) &= C + \Omega_c y(t) \\ \frac{dv_y}{dt} &= -\frac{qB}{\gamma m} (C + \Omega_c y(t)) \\ \frac{d^2 y}{dt^2} + \Omega_c^2 y &= -\frac{qB}{\gamma m} C \\ y(t) &= -\frac{C}{\Omega_c} + A' \cos(\Omega_c t + \delta) \\ v_x(t) &= \Omega_c A' \cos(\Omega_c t + \delta) \quad v_y(t) = -\Omega_c A' \sin(\Omega_c t + \delta) \quad \therefore \Omega_c A' = v_0 \\ x(t) &= C' + A' \sin(\Omega_c t + \delta) \\ A' &= r = \frac{v_0}{\Omega_c}\end{aligned}$$

4. This problem is a perfect example of the use of the magnetic rigidity. The electron relativistic momentum is  $p = \sqrt{\gamma^2 - 1} m_0 c = \sqrt{\gamma^2 - 1} (0.511 \text{ MeV}/c)$ , the magnetic rigidity is  $\sqrt{\gamma^2 - 1} (0.511 \text{ MV}/c)$ , and  $1 \text{ T} = 1 \text{ (V sec)/(m}^2)$ . Now

$$B = \frac{(B\rho)}{L} 2 \sin(\theta/2)$$

for magnets in the normal configuration. The bend angles are  $\pi/16 = 0.19635$  rad for the first arc and  $\pi/32 = 0.098175$  rad for the rest of the arcs.

$$B_1 = \frac{\sqrt{(605/0.511)^2 - 1} \times 0.511 \times 10^6 \text{ V}}{2.998 \times 10^8 \text{ m/sec}(1 \text{ m})} 2 \sin 0.098175 = 0.3956 \text{ T} = 3.956 \text{ kG}$$

$$B_2 = \frac{\sqrt{(1693/0.511)^2 - 1} \times 0.511 \times 10^6 \text{ V}}{2.998 \times 10^8 \text{ m/sec}(1 \text{ m})} 2 \sin 0.04909 = 0.5542 \text{ T} = 5.542 \text{ kG}$$

$$B_3 = \frac{\sqrt{(2781/0.511)^2 - 1} \times 0.511 \times 10^6 \text{ V}}{2.998 \times 10^8 \text{ m/sec}(2 \text{ m})} 2 \sin 0.04909 = 0.4552 \text{ T} = 4.552 \text{ kG}$$

$$B_4 = \frac{\sqrt{(3868/0.511)^2 - 1} \times 0.511 \times 10^6 \text{ V}}{2.998 \times 10^8 \text{ m/sec}(3 \text{ m})} 2 \sin 0.04909 = 0.4220 \text{ T} = 4.220 \text{ kG}$$

$$B_5 = \frac{\sqrt{(4956/0.511)^2 - 1} \times 0.511 \times 10^6 \text{ V}}{2.998 \times 10^8 \text{ m/sec}(3 \text{ m})} 2 \sin 0.04909 = 0.5408 \text{ T} = 5.408 \text{ kG}$$

Arc	Electron Energy (MeV)	Number of Dipoles	Dipole Length (m)	Bend Angle (rad)	Magnetic Field (T)
1	605	16	1	0.19635	0.3956
2	1693	32	1	0.098175	0.5542
3	2781	32	2	0.098175	0.4552
4	3868	32	3	0.098175	0.4220
5	4956	32	3	0.098175	0.5408

5. If  $\phi = 180^\circ$ , then the transfer matrix is

$$\begin{pmatrix} -1 & 0 \\ 0 & -1 \end{pmatrix}$$

In practice this condition may be achieved with a single transfer matrix, 2 transfer matrices of 90 degrees phase advance, 3 of 60 degrees of phase advance, etc.