

Physics 417/517

Introduction to Particle Accelerator Physics

G. A. Krafft

Jefferson Lab

Jefferson Lab Professor of Physics

Old Dominion University

Synchrotron Radiation



Accelerated particles emit electromagnetic radiation. Emission from very high energy particles has unique properties for a radiation source. As such radiation was first observed at one of the earliest electron synchrotrons, radiation from high energy particles (mainly electrons) is known generically as synchrotron radiation by the accelerator and HENP communities.

The radiation is highly collimated in the beam direction

From relativity

$$ct' = \gamma ct - \gamma\beta z$$

$$x' = x$$

$$y' = y$$

$$z' = -\gamma\beta ct + \gamma z$$

Lorentz invariance of wave phase implies $k^\mu = (\omega/c, k_x, k_y, k_z)$ is a Lorentz 4-vector

$$\omega' = \gamma\omega - \gamma\beta k_z c$$

$$k'_x = k_x$$

$$k'_y = k_y$$

$$k'_z = -\gamma\beta c\omega + \gamma k_z$$

$$\sin \theta = \frac{\sqrt{k_x^2 + k_y^2}}{\omega / c} \quad \sin \theta' = \frac{\sqrt{k_x'^2 + k_y'^2}}{\omega' / c} \quad \cos \theta' = \frac{k'_z}{\omega' / c}$$

$$\omega / c = \gamma\omega' / c + \gamma\beta k'_z = \gamma (1 + \beta \cos \theta') (\omega' / c)$$

$$\theta \approx \sin \theta = \frac{\sin \theta'}{\gamma (1 + \beta \cos \theta')}$$

Therefore all radiation with $\theta' < \pi / 2$, which is roughly $1/2$ of the emission for dipole emission from a transverse acceleration in the beam frame, is Lorentz transformed into an angle less than $1/\gamma$. Because of the strong Doppler shift of the photon energy, higher for $\theta \rightarrow 0$, most of the energy in the photons is within a cone of angular extent $1/\gamma$ around the beam direction.

Larmor's Formula



For a particle executing non-relativistic motion, the total power emitted in electromagnetic radiation is (Larmor)

$$P(t) = \frac{1}{6\pi\epsilon_0} \frac{q^2}{c^3} |\vec{a}|^2 = \frac{1}{6\pi\epsilon_0} \frac{e^2}{m^2 c^3} |\dot{\vec{p}}|^2$$

Lienard's relativistic generalization: Note both dE and dt are the fourth component of relativistic 4-vectors when one is dealing with photon emission. Therefore, their ratio must be an Lorentz invariant. The invariant that reduces to Larmor's formula in the non-relativistic limit is

$$P = - \frac{e^2}{6\pi\epsilon_0 c} \frac{du^\mu}{d\tau} \frac{du_\mu}{d\tau}$$

$$P(t) = \frac{e^2}{6\pi\epsilon_0 c} \gamma^6 \left(\dot{\vec{\beta}}^2 - \left[\vec{\beta} \times \dot{\vec{\beta}} \right]^2 \right)$$

For acceleration along a line, second term is zero and first term for the radiation reaction is small compared to the acceleration as long as gradient less than 10^{14} MV/m. Technically impossible.

For transverse bend acceleration $\dot{\vec{\beta}} = -\frac{\beta^2 c}{\rho} \hat{r}$

$$P(t) = \frac{e^2 c}{6\pi\epsilon_0 \rho^2} \beta^4 \gamma^4$$

Fractional Energy Loss



$$\delta E = \frac{e^2}{6\pi\epsilon_0\rho} \Theta \beta^3 \gamma^4$$

For one turn with isomagnetic bending fields

$$\frac{\delta E}{E_{beam}} = \frac{4\pi r_e}{3\rho} \beta^3 \gamma^3$$

r_e is the classical electron radius: 2.82×10^{-13} cm

Radiation Power Distribution



Consulting your favorite Classical E&M text (Jackson, Schwinger, Landau and Lifshitz Classical Theory of Fields)

$$\frac{dP}{d\omega} = \frac{\sqrt{3}}{8\pi^2 \epsilon_0} \frac{e^2}{\rho} \gamma \frac{\omega}{\omega_c} \int_{\omega/\omega_c}^{\infty} K_{5/3}(x) dx$$

Critical Frequency

Critical (angular) frequency is

$$\omega_c = \frac{3}{2} \gamma^3 \frac{c}{\rho}$$

Energy scaling of critical frequency is understood from $1/\gamma$ emission cone and fact that $1 - \beta \sim 1/(2\gamma^2)$

$$t_A = -\frac{\rho}{\gamma\beta c}$$

$$t_B = \frac{\rho}{\gamma\beta c} \approx \frac{\rho}{\gamma c} + \frac{\rho}{2\gamma^3 c}$$

$$\Delta t = t_B - \left(\frac{\rho}{\gamma c} - \frac{\rho}{2\gamma^3 c} \right) \approx \frac{\rho}{\gamma^3 c}$$

